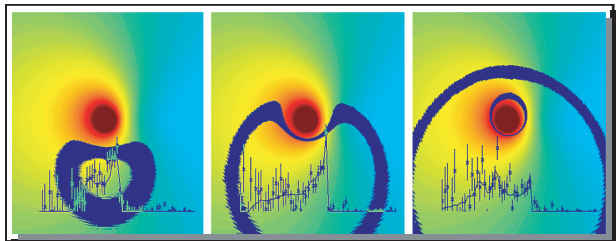


**Constellation-X will use X-ray Spectroscopy to revolutionize our knowledge of the universe. It will probe closer to black hole event horizons with 100 times better sensitivity than ever before.**

Constellation-X's four science objectives are tightly connected to NASA SEU themes:



- I. Measure the effects of **strong gravity** near the event horizon of supermassive black holes.  
*What is the nature of space and time?*  
*What powers supermassive black holes?*
- II. Trace visible matter throughout the universe and constrain the nature of **dark matter** and dark energy.  
*What is the universe made of?*  
*How does the universe evolve?*
- III. Study the formation of supermassive **black holes** and trace their evolution with cosmic time.  
*What roles do they play in galaxy evolution?*  
*What is the total energy output of the universe?*
- IV. Study the **life cycles of matter** and energy and understand the behavior of matter in extreme environments.  
*What new forms of matter will be discovered?*  
*How does the chemical composition of the universe evolve?*

**Well-defined science objectives provide well-defined measurement requirements.**

**Management: A straightforward approach with few interfaces and highly experienced teams:**

- Mission managed by NASA/GSFC
- SAO part of management team
- Prime contractor for observatory

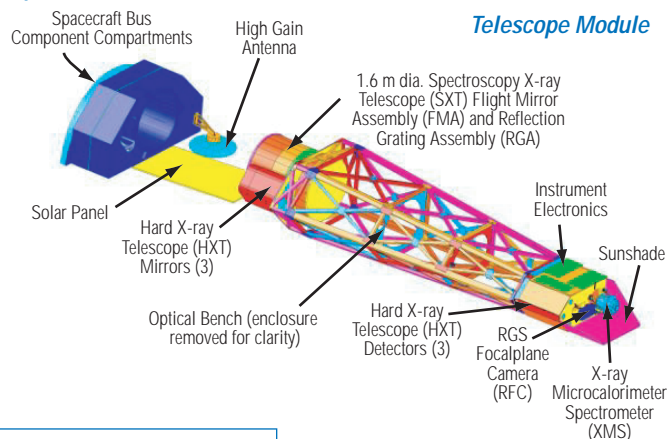
### Mission Overview:

Launch dates:	2010 and 2011
Launch vehicles:	Atlas V (two)
Constellation:	4 observatories point at target (no formation flying)
Mission lifetime:	4 years for fully operational constellation & 10 year goal
Orbit:	L2 Lissajous
NASA mission cost:	\$1,597B (RY)

### Exploded view of a Constellation-X observatory

- Robust, modular mission design
- Performance verifiable on the ground
- Meets mission requirements traceable to the science objectives

#### Spacecraft Bus



### Constellation-X provides:

High observing efficiency (90%)  
 Large sample sizes of key astrophysical objects  
 Broad-band X-ray imaging spectroscopy (0.25 - 40 keV)  
 General observer facility with programs selected by peer review to carry out world-class science  
 Dramatic improvements in spectroscopic sensitivity, about a factor of 100 over previous missions

## Observatory Characteristics:

Number: 4 identical observatories  
Wet mass (each): ~2480 kg  
Power (each): ~1075 W  
Data storage: 42 Gbit  
Source location: 5 arcsec, post facto  
Attitude control: 3-axis stabilized  
~30 arcsec control  
Communication: X/S-band downlink  
(1.7 Mbps/2 kbps)  
S-band uplink (2 kbps)

Mechanisms: Few; low precision focusing

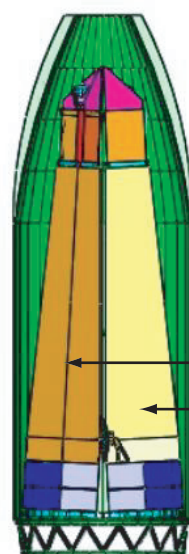
**No new spacecraft technologies are required**

Modular design minimizes interfaces and simplifies I&T flow

## Key Heritage Elements:

- Technologies evolve from existing hardware.
- Our teams bring significant flight experience.
- Chandra provides heritage for systems engineering, key mechanisms, I&T, and the Science and Operations Center.

Technology Area	Current TRL	Date for TRL6	Heritage
Mirrors (SXT, HXT)	3/4	FY06	ASCA, Astro-E2, InFocus Chandra, XMM-Newton
XMS Microcalorimeter	4	FY05	Astro-E2
HXT detector	4/6	FY05	HEFT, InFocus
RGS CCDs	3	FY05	ASCA, Chandra
RGS gratings	3	FY06	XMM-Newton
XMS Cryocooler	4	FY06	HST, TES, AIRS
XMS ADR	4	FY06	Astro-E2



**Resource margins are based on a mature mission concept**

Mass Margin: 34%  
Power Margin: 34%  
Schedule Contingency: 5 months (10%)  
(plus slack)  
Cost Reserves: \$191M (22%)  
(observatory development)

**Observatory 1**

**Observatory 2**

Two observatories are packaged inside an Atlas V fairing

**Science Payload: Instruments are extensions of recent, flight-proven hardware, minimizing technology development risks while meeting requirements with adequate performance margins.**

SXT FMA: Primary optic feed for XMS and RGS  
 $\approx 15,000 \text{ cm}^2$  at 1.25 keV

RGS: Dispersive spectrometer from 0.25 - 2 keV  
Resolving power  $R \approx 300$  at 0.6 keV

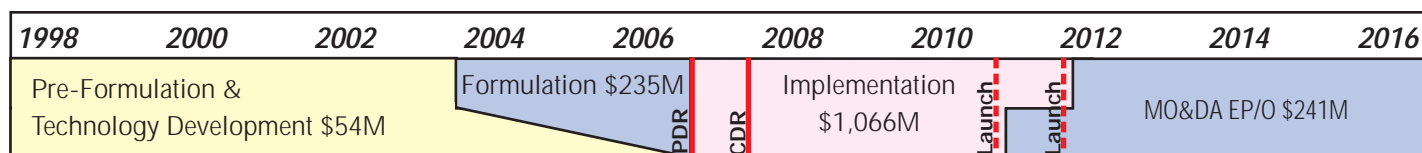
XMS: Imaging spectrometer from 0.6 - 10 keV  
Resolving power  $R \approx 1500$  at 6.0 keV

HXT: Imaging spectrometer from 6 - 40 keV  
Resolving power  $R \approx 10$  at 40 keV

## Why Constellation-X Now?

- Guaranteed, compelling science returns
- Breakthrough discoveries require comprehensive spectroscopic studies
- Addresses priorities of the NASA SEU program
- Technology development has demonstrated readiness to proceed; team is in place

## Constellation-X Schedule



RY DOLLARS